

## Solutions of Einstein's Equations: The Field Outside Of A Massive Charged Spherical Body

Below, we list the non-zero components of all of the relevant tensors and other objects which are involved in showing that the Reissner–Nordstrom solution ( $Q \neq 0$ ) and the Schwarzschild solution ( $Q = 0$ ) are solutions to Einstein's equations.

### The (covariant) Metric

$$\begin{aligned}g_{tt} &= -\left(1 - 2\frac{m}{r} + \frac{Q^2}{r^2}\right) \\g_{rr} &= \left(1 - 2\frac{m}{r} + \frac{Q^2}{r^2}\right)^{-1} \\g_{\theta\theta} &= r^2 \\g_{\phi\phi} &= r^2 \sin^2\theta\end{aligned}$$

### The (contravariant) Inverse Metric

$$\begin{aligned}g^{tt} &= -\left(1 - 2\frac{m}{r} + \frac{Q^2}{r^2}\right)^{-1} \\g^{rr} &= \left(1 - 2\frac{m}{r} + \frac{Q^2}{r^2}\right) \\g^{\theta\theta} &= \frac{1}{r^2} \\g^{\phi\phi} &= \frac{1}{r^2 \sin^2\theta}\end{aligned}$$

## The Affine (metric) Connection

$$\begin{aligned}
 \Gamma_{tr}^t &= \frac{m r - Q^2}{r (r^2 - 2 m r + Q^2)} \\
 \Gamma_{tt}^r &= \frac{(r^2 - 2 m r + Q^2) (m r - Q^2)}{r^5} \\
 \Gamma_{rr}^r &= -\frac{m r - Q^2}{r (r^2 - 2 m r + Q^2)} \\
 \Gamma_{\theta\theta}^r &= -\frac{r^2 - 2 m r + Q^2}{r} \\
 \Gamma_{\phi\phi}^r &= -\frac{(r^2 - 2 m r + Q^2) \sin^2\theta}{r} \\
 \Gamma_{\phi\theta}^\theta &= \frac{1}{r} \\
 \Gamma_{\phi\phi}^\theta &= -\sin\theta \cos\theta \\
 \Gamma_{r\phi}^\phi &= \frac{1}{r} \\
 \Gamma_{\theta\phi}^\phi &= \frac{\cos\theta}{\sin\theta}
 \end{aligned}$$

## The Riemann–Christoffel Tensor

$$\begin{aligned}
 R_{trtr} &= \frac{2 m r - 3 Q^2}{r^4} \\
 R_{t\theta t\theta} &= -\frac{(r^2 - 2 m r + Q^2) (m r - Q^2)}{r^4} \\
 R_{t\phi t\phi} &= -\frac{(r^2 - 2 m r + Q^2) (m r - Q^2) \sin(\theta)^2}{r^4} \\
 R_{r\theta r\theta} &= \frac{m r - Q^2}{r^2 - 2 m r + Q^2} \\
 R_{r\phi r\phi} &= \frac{(m r - Q^2) \sin^2\theta}{r^2 - 2 m r + Q^2} \\
 R_{\theta\phi\theta\phi} &= -2 \sin^2\theta m r + \sin^2\theta Q^2
 \end{aligned}$$

### The Ricci tensor

$$\begin{aligned}R_{tt} &= \frac{(r^2 - 2 m r + Q^2) Q^2}{r^6} \\R_{rr} &= -\frac{Q^2}{r^2 (r^2 - 2 m r + Q^2)} \\R_{\theta\theta} &= \frac{Q^2}{r^2} \\R_{\phi\phi} &= \frac{\sin^2\theta Q^2}{r^2}\end{aligned}$$

### The Ricci Scalar

$$R = 0$$

**The Einstein Tensor:**  $G_{\mu\nu} \equiv R_{\mu\nu} - \frac{1}{2}g_{\mu\nu}R$ .

$$\begin{aligned}G_{tt} &= \frac{(r^2 - 2 m r + Q^2) Q^2}{r^6} \\G_{rr} &= -\frac{Q^2}{r^2 (r^2 - 2 m r + Q^2)} \\G_{\theta\theta} &= \frac{Q^2}{r^2} \\G_{\phi\phi} &= \frac{\sin^2\theta Q^2}{r^2}\end{aligned}$$

## The Maxwell Tensor

$$F_{tr} = E_r(r) = -F_{rt}.$$

**The Energy–Momentum Tensor:**  $4\pi T_{\mu\nu} = -g^{\alpha\beta} F_{\mu\alpha} F_{\nu\beta} + \frac{1}{4} g_{\mu\nu} F_{\alpha\beta} F^{\alpha\beta}$

$$\begin{aligned} T_{tt} &= -\frac{E_r(r)^2(r^2 - 2mr + Q^2)}{2r^2} \\ T_{rr} &= \frac{E_r(r)^2 r^2}{2(r^2 - 2mr + Q^2)} \\ T_{\theta\theta} &= -\frac{E_r(r)^2 r^2}{2} \\ T_{\phi\phi} &= -\frac{E_r(r)^2 r^2 \sin^2\theta}{2} \end{aligned}$$

So we see that Einstein's equations:

$$R_{\mu\nu} - \frac{1}{2} g_{\mu\nu} R = -8\pi T_{\mu\nu}$$

are satisfied with the field for a point charge  $E_r(r) = Q/r^2$ .

For  $Q=0$  we have the field/geometry outside a neutral spherical body of mass  $m$ . The source terms vanish identically. This is an 'empty-space' solution of Einstein's equations. For the  $Q \neq 0$  case, we have the field/geometry outside a spherical body of electric charge  $Q$  and mass  $m$ . It is not an empty-space solution. Space is filled with the 'field lines' due to the point source of charge at  $r = 0$ .

You have studied many properties of these solutions in class, exercises and in the midterm exam.

–Clifford V. Johnson